

On Two-Field Solitons in 2 and 3 Dimensions

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Introduction

- Theoretical description of experimental results obtained at present-day accelerators requires **nonperturbative** quantum study of the Standard Model (**SM**)
 - **SM=QCD+EW**[Glashow, Weinberg, Salam]
 - **Nonperturbative QFTs. Approaches available:**
 - i) Lattice numerical simulations - from first principles
 - ii) Semianalytical techniques, e.g. Dyson-Schwinger (drawback: ambiguities)
 - iii) Quantization using **localized classical solutions** (instantons, dyons, **solitons**)
- ⇒ The main problem of solitonic approach - to find **stable 3D** solitons in realistic FTs.
Yet **not solved**, but **not forbidden** by theorems, and **there is hope...**

Gauged NSMs vs EW models

We study 2-field gauge-invariant **NSMs** (nonlinear sigma models) whose structure resembles that of bosonic sector of **EW** theory.

In fact, **our gauged NSMs** include:

i) unit length scalar N -component field, with values on S^{N-1} :

$$s_1^2 + \dots s_N^2 = 1 \quad (N = 3, 4)$$

interacting with

ii) vector field with $U(1)$ or $SU(2)$ symmetry (Maxwell or Yang-Mills).

Compare with **EW bosonic sector** -

omitting there $U(1)$ field and freezing radial component of Higgs field one gets **$SU(2)$ -Higgs frozen model**. It contains:

i) unit 4-component scalar field with values on S^3

ii) vector $SU(2)$ Yang-Mills field.

⇒ **Clear resemblance !**

⇒ Both gauged NSMs and EW theory are invariant with respect to local gauge transformations.

A3M model in 2 dimensions

$$\mathcal{L} = \bar{\mathcal{D}}_\mu s_- \mathcal{D}^\mu s_+ + \partial_\mu s_3 \partial^\mu s_3 - V(s_a) - \frac{1}{4} F_{\mu\nu}^2,$$

$$\bar{\mathcal{D}}_\mu = \partial_\mu + ieA_\mu, \quad \mathcal{D}_\mu = \partial_\mu - ieA_\mu,$$

$$s_+ = s_1 + is_2, \quad s_- = s_1 - is_2,$$

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu, \quad V(s_a) = \beta(1 - s_3^2),$$

where β, e are coupling constants and $\mu, \nu = 0, 1, 2$. We look for the topological solitons of the A3M model using the “hedgehog-like” ansatz for the A3-field

$$s_1 = \cos m\chi \sin \theta(R), \quad s_2 = \sin m\chi \sin \theta(R),$$

$$s_3 = \cos \theta(R),$$

$$\sin \chi = \frac{y}{R}, \quad \cos \chi = \frac{x}{R}, \quad R^2 = x^2 + y^2,$$

where m is an integer number, and the standard “vortex” ansatz for the vector field A_μ , describing localized distributions of a stationary magnetic field:

$$A_0 = 0, \quad A_1 = A_x = -ma(R) \frac{y}{R^2},$$

$$A_2 = A_y = ma(R) \frac{x}{R^2}.$$

Topological charge (mapping degree $R_{comp}^2 \rightarrow S^2$)

$$Q_t = m$$

Making rescaling

$$a = \alpha e^{-1}, \quad R = r e^{-1},$$

A3M (2)

We get for stationary Hamiltonian density $\mathcal{H}(r) = e^{-2}\mathcal{H}(R)$:

$$\mathcal{H}(r) = \left(\frac{d\theta}{dr}\right)^2 + \sin^2\theta \left[p + \frac{m^2(\alpha - 1)^2}{r^2} \right] + \frac{m^2}{2} \left(\frac{1}{r} \frac{d\alpha}{dr} \right)^2,$$

$$H = \int_0^\infty \mathcal{H}(r) 2\pi r dr, \quad p = \frac{\beta}{e^2}.$$

Calculating $\delta H/\delta\theta$ and $\delta H/\delta\alpha$, we obtain coupled equations for $\theta(r)$ and $\alpha(r)$:

$$\frac{d^2\theta}{dr^2} + \frac{1}{r} \frac{d\theta}{dr} - \sin\theta \cos\theta \left[\frac{m^2(\alpha - 1)^2}{r^2} + p \right] = 0,$$

$$\frac{d^2\alpha}{dr^2} - \frac{1}{r} \frac{d\alpha}{dr} + 2\sin^2\theta(1 - \alpha) = 0,$$

to be solved under the following boundary conditions:

$$\begin{aligned} \theta(0) &= \pi, & \theta(\infty) &= 0, \\ \alpha(0) &= 0, & \frac{d\alpha}{dr}(\infty) &= 0. \end{aligned}$$

A3M (3)

Using series expansion of $\theta(r)$ and $\alpha(r)$ at $r \rightarrow 0$, we find from Eqs. (20) and (21) for $m = 1$

$$\theta(r) = \pi - C_1 r + o(r),$$

$$\alpha(r) = r^2 \left(E_1^2 - \frac{1}{4} C_1^2 r^2 \right) + o(r^4).$$

and for $m = 2$

$$\theta(r) = \pi - C_2 r^2 + o(r^2),$$

$$\alpha(r) = r^2 \left(E_2^2 - \frac{1}{12} C_2^2 r^4 \right) + o(r^6).$$

The asymptotic form of the soliton solution for $r \rightarrow \infty$ is:

$$\theta(r) \approx \frac{T}{\sqrt{r}} \exp(-\sqrt{pr}), \quad T = \text{const},$$

$$\alpha(r) \approx \alpha_\infty - (1 - \alpha_\infty) \frac{T^2}{2rp} \exp(-2\sqrt{pr}).$$

Solutions exist for $0 < p < p_{cr} \approx 0.41$.

A3M: solutions(4)

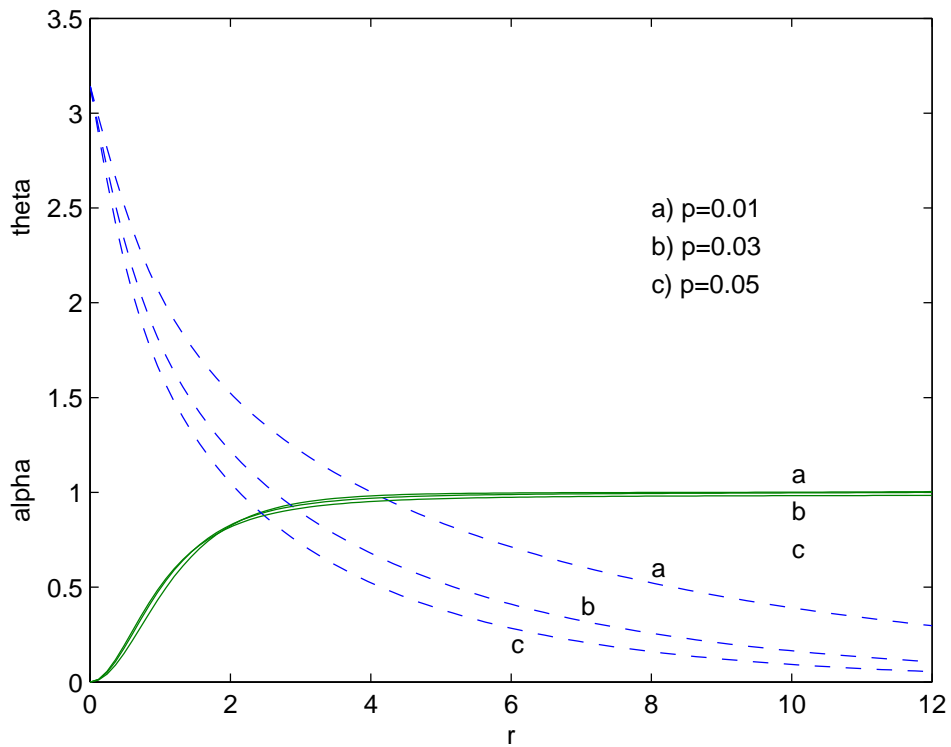


Figure 1: $\theta(r)$ and $\alpha(r)$ for $p = 0.01, 0.03, 0.05$.

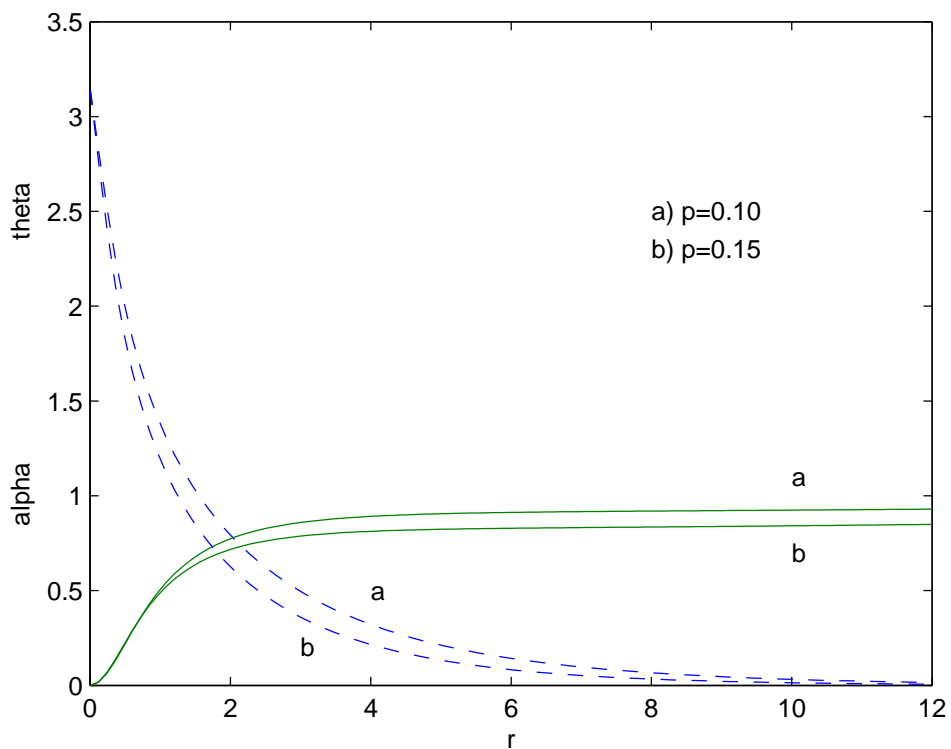


Figure 2: $\theta(r)$ and $\alpha(r)$ for $p = 0.1, 0.15$.

A3M: energy density and $B(r)$

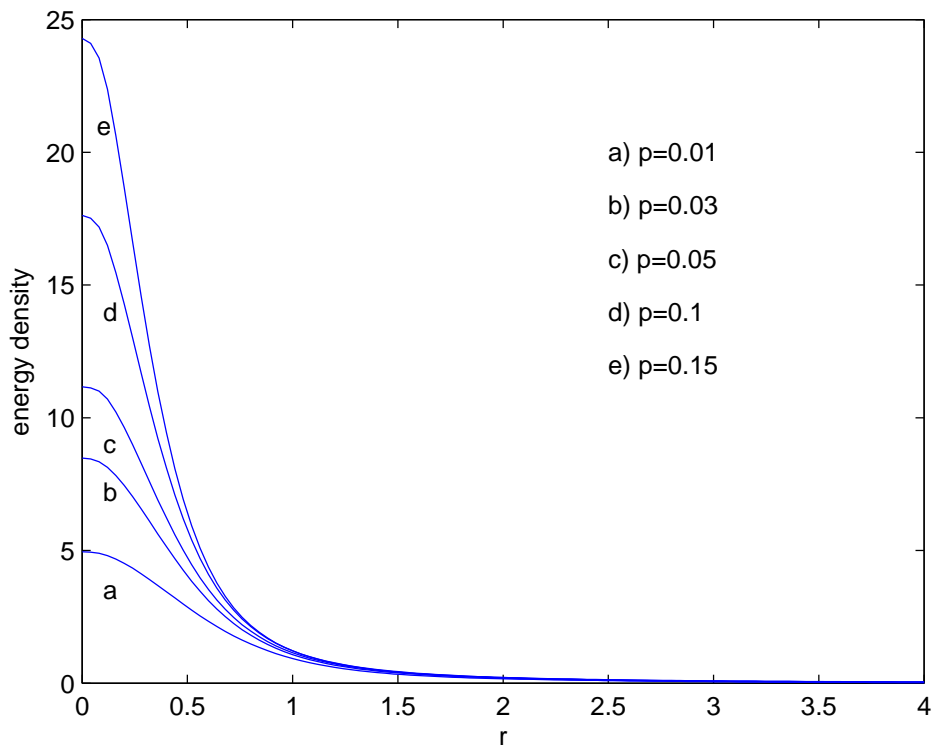


Figure 3: Energy density for various p .

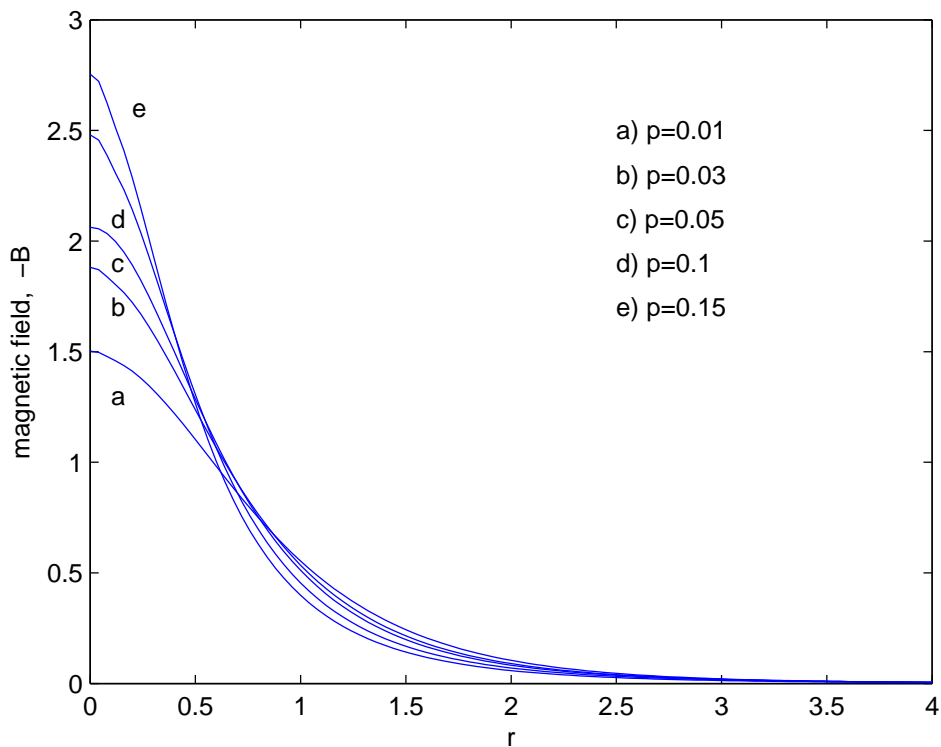


Figure 4: Magnetic field for various p .

A3M: Soliton energy and $\alpha(\infty)$ vs p

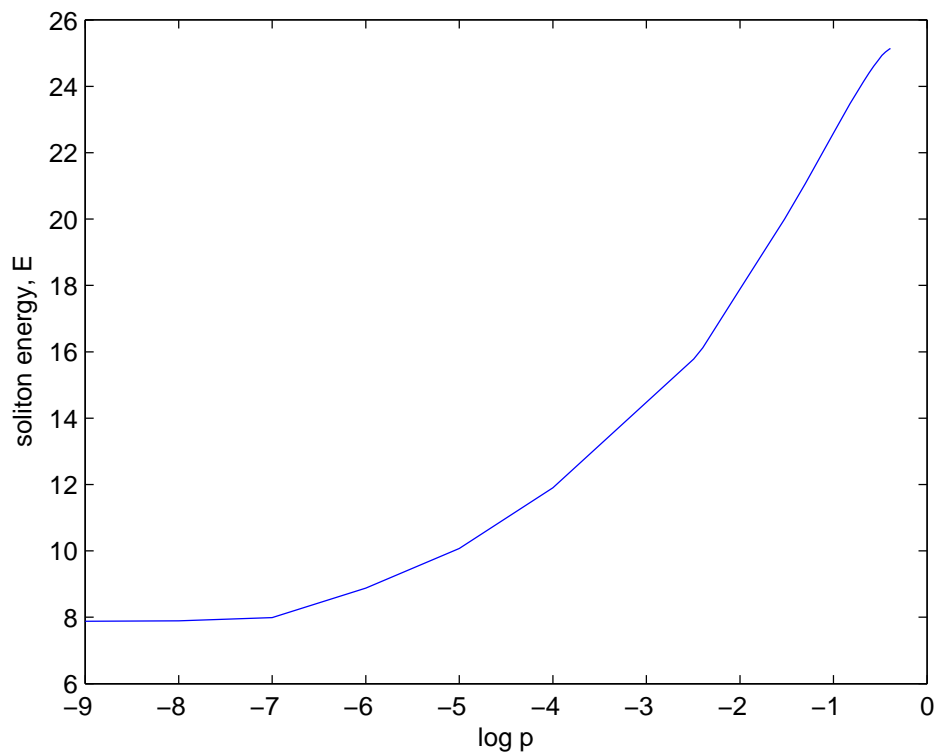


Figure 5: Soliton energy for various p .

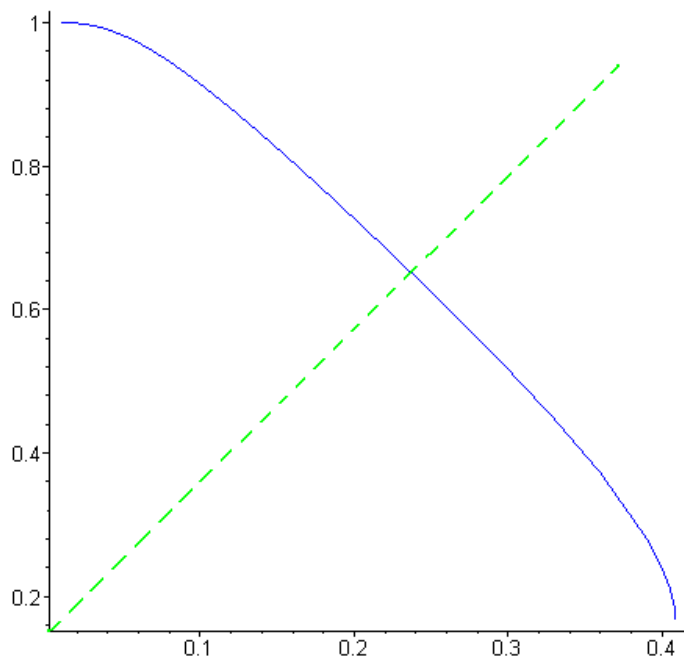


Figure 6: $\alpha(\infty)$ vs p .

$Q_t = 2$ solitons: $m = 2$.

Numerically found, that

$E_{sol}(Q_t = 2, p) < 2 * E_{sol}(Q_t = 1, p)$ for all $0 < p < p_{cr}$,

hence two solitons attract to each other.

CONCLUSIONS

\Rightarrow For $p < p_{cr}$ A3M solitons do exist and prove to be stable

\Rightarrow One can quantize the model around these solitons, quasiclassically (WKB) or in lattice approach.

\Rightarrow One can consider 2D A3M solitons as extended strings embedded into 3D space (e.g., “cosmic strings”).

A4YM model

(1)

- Includes unit 4-component scalar field $q^\alpha(x^\mu)$ (“the A4-field”) interacting with the vector $SU(2)$ Yang-Mills field $A_\mu^a(x^\nu)$

$$\mathcal{L} = \mathcal{D}_\mu q^a \mathcal{D}^\mu q^a + \partial_\mu q^0 \partial^\mu q^0 - V(q^0) - \frac{1}{4} (F_{\mu\nu}^a)^2,$$

$$\mathcal{D}_\mu q^a = \partial_\mu q^a + g \varepsilon^{abc} A_\mu^b q^c,$$

$$F_{\mu\nu}^a = \partial_\mu A_\nu^a - \partial_\nu A_\mu^a + g \varepsilon^{abc} A_\mu^b A_\nu^c,$$

$$V(q^0) = \beta [1 - (q^0)^2],$$

$$\alpha, \mu, \nu = 0, 1, 2, 3, \quad a, b, c = 1, 2, 3, \quad \beta, g - \text{const.}$$

- We use ansatzes:

$$q^0 = \cos \theta(R), \quad q^a = \sin \theta(R) \frac{x^a}{R}, \quad R^2 = x^2 + y^2 + z^2,$$

$$A_0^a = 0, \quad A_i^a = c(R) \varepsilon^{iak} x^k.$$

- Then the Hamiltonian density distributions of localized field bunches are spherically symmetric:

$$\mathcal{H}_{st}(R) = \left(\frac{d\theta}{dR} \right)^2 + \frac{2\sin^2\theta}{R^2} + 4gc\sin^2\theta + 2g^2c^2R^2\sin^2\theta$$

$$+ 6c^2 + \left(\frac{dc}{dR} \right)^2 R^2 + \frac{1}{2}g^2c^4R^4 + 4Rc \frac{dc}{dR} + 2gR^2c^3 + \beta\sin^2\theta.$$

A4YM model (2)

Introduce dimensionless variables

$$r = gR, \quad b(r) = g^{-1}cr^2.$$

Calculating $\delta\mathcal{H}/\delta\theta$ and $\delta\mathcal{H}/\delta b$, we get coupled equations ($P = \frac{\beta}{g^2}$)

$$\frac{d^2\theta}{dr^2} + \frac{2d\theta}{rdr} - \sin\theta\cos\theta \left[\frac{2(b+1)^2}{r^2} + P \right] = 0,$$

$$\frac{d^2b}{dr^2} - \frac{2b}{r^2} - 2\sin^2\theta(1+b) - \frac{b^2}{r^2}(b+3) = 0.$$

We looked for localized solutions, setting the following boundary conditions:

$$\theta(0) = \pi, \quad \theta(\infty) = 0,$$

$$b(0) = 0, \quad b(\infty) = B.$$

Solutions to above problem would define localized distributions $q^\alpha(x^k)$, $\alpha = 0, 1, 2, 3$, $k = 1, 2, 3$, of the A4-field with unit topological charge, $Q_t = 1$. Here Q_t is the "mapping degree", of continuous maps $R_{comp}^3 \rightarrow S^3$. However such solutions have not been found. Because of that we look for more general ansatz.

A4YM model

(3)

More general ansatz keeps hedgehog form for q^α and generalized ansatz for A_i :

$$A_i^a(x) = \epsilon_{aij} n_j C(r) r + (\delta_{ai} - n_a n_i) \frac{B(r)}{r} + n_a n_i \frac{G(r)}{r}.$$

However such ansatz should respect Coulomb (Lorentz) gauge. Equating $\delta A_i^a / \delta x_i = 0$, we found $C(r) = B(r) + \text{const}$. And finally we choose the ansatz

$$A_i^a(x) = \epsilon_{aij} n_j C(r) r + \delta_{ai} \frac{B(r)}{r} + \frac{G n_a n_i}{r}.$$

⇒ Now calculate hamiltonian density for such ansatz.

Equating variation derivatives $\frac{\delta \mathcal{H}(r)}{\delta C}$, $\frac{\delta \mathcal{H}(r)}{\delta B}$, $\frac{\delta \mathcal{H}(r)}{\delta \theta}$ to 0, we obtain coupled equations for radial functions $C(r)$, $B(r)$, $\theta(r)$. Their solution should define localized (solitonic) solutions to A4YM model.

⇒ The study is in progress.

$$\begin{aligned}
Hst := & \frac{1}{2} R^4 g^2 c(R)^4 + \left(-2 g c(R)^3 + 2 g^2 c(R)^2 - 2 c(R)^2 g^2 \cos(\theta)^2 + \left(\frac{d}{dR} c(R) \right)^2 \right) R^2 \\
& + 4 R \left(\frac{d}{dR} c(R) \right) c(R) + 6 c(R)^2 + 9 g^2 c(R)^2 G^2 + 18 g^2 B(R)^2 c(R)^2 + 18 B(R) G g^2 c(R)^2 \\
& + \left(\frac{d}{dR} \theta(R) \right)^2 - 4 g c(R) + 4 g c(R) \cos(\theta)^2 \\
& - \frac{18 g \left(B(R) \left(\frac{d}{dR} c(R) \right) - \left(\frac{d}{dR} B(R) \right) c(R) \right) (B(R) + G)}{R} + \left(-18 g^2 B(R)^2 \cos(\theta)^2 \right. \\
& + 9 \left(\frac{d}{dR} B(R) \right)^2 + 2 + 18 g^2 B(R)^2 - 2 \cos(\theta)^2 - 18 g B(R)^2 c - 18 g c(R) G^2 \\
& \left. - 54 g B(R)^2 c(R) - 72 B(R) G g c(R) \right) / R^2 - \frac{18 \left(\frac{d}{dR} B(R) \right) (B(R) + G)}{R^3} \\
& + \frac{162 G B(R)^3 g^2 + 9 G^2 + 81 B(R)^2 G^2 g^2 + 18 G B(R) + \frac{243}{2} B(R)^4 g^2 + 9 B(R)^2}{R^4}
\end{aligned}$$